

## SWOT CalVal Planning Meeting

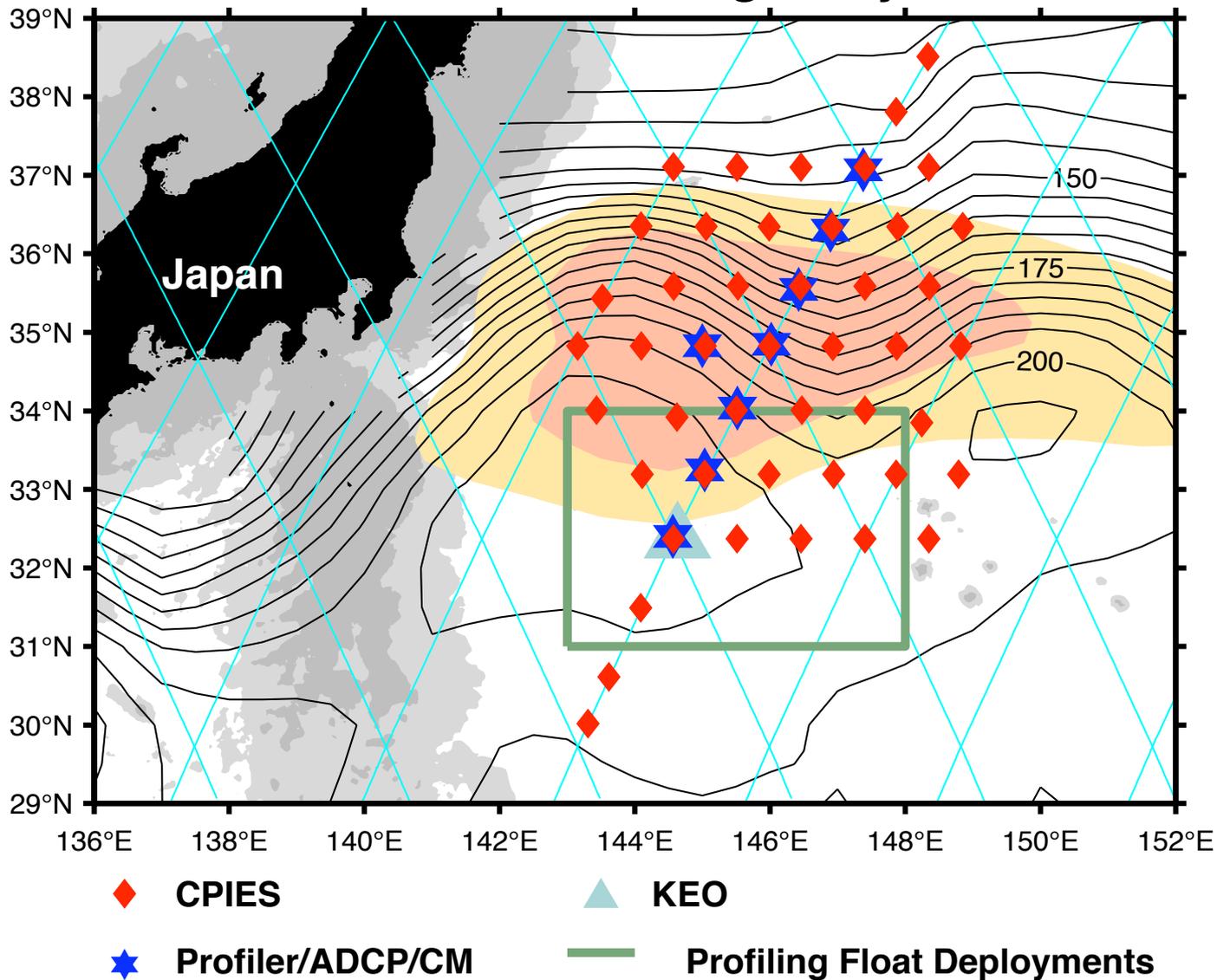
NASA-JPL, Pasadena, CA -- 2016-Jun-16

*D. R. Watts, K. Donohue (URI) & M. Andres (WHOI)*

- PIES –
  - What are they?
  - How do we use them to get SSH?
- PIES\_SSH uncertainty –
  - Where does it come from?
  - What measurements will reduce uncertainty?

# Kuroshio Extension System Study (KESS, 6/2004- 6/2006)

## KESS Observing Array



# PIES

(Pressure-recording Inverted Echo Sounder)



- Emits 12 kHz sound pulses
- Measures  $\tau$  = round-trip travel time from bottom to surface
- Measures bottom pressure ( $P_{\text{bot}}$ ) – resolution < 1 mm
- A robust empirical relationship exists between  $\tau$  and vertical profiles  $T(P)$ ,  $S(P)$ ,  $\delta(P)$ 
  - ➔ geopotential  $\Phi$

## SSH COMPRISES STERIC AND MASS-LOAD PARTS

Sea Surface Height ( $\eta$ ) after IB compensation:

$$(\eta' - IB) = \Phi'/g + P'/\rho g$$

Where  $(\bullet)'$  = anomaly from mean

$\Phi$  = Geopotential Height

(0 re near-bottom nominal isobar, like 4000 or 5000 dbar)

$g$  = Gravity

$P$  = Bottom Pressure

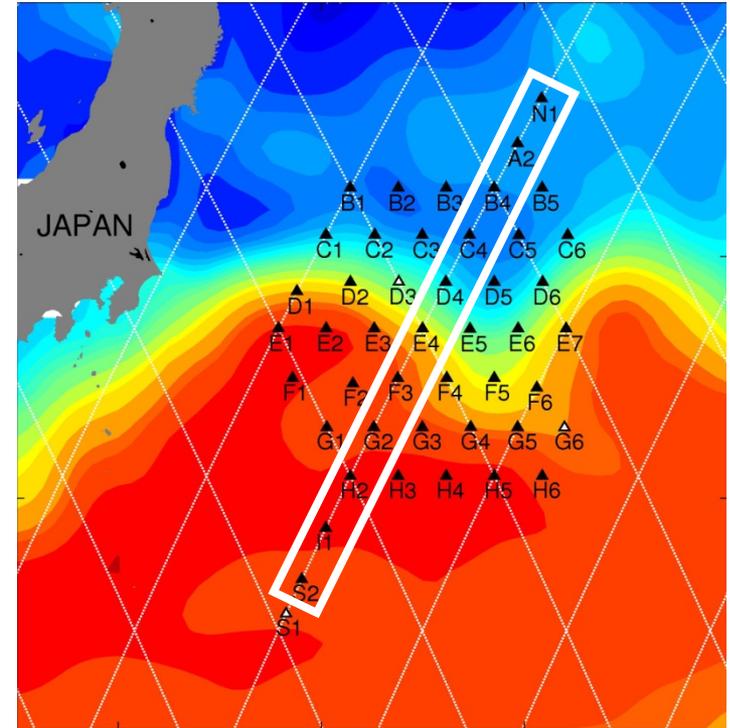
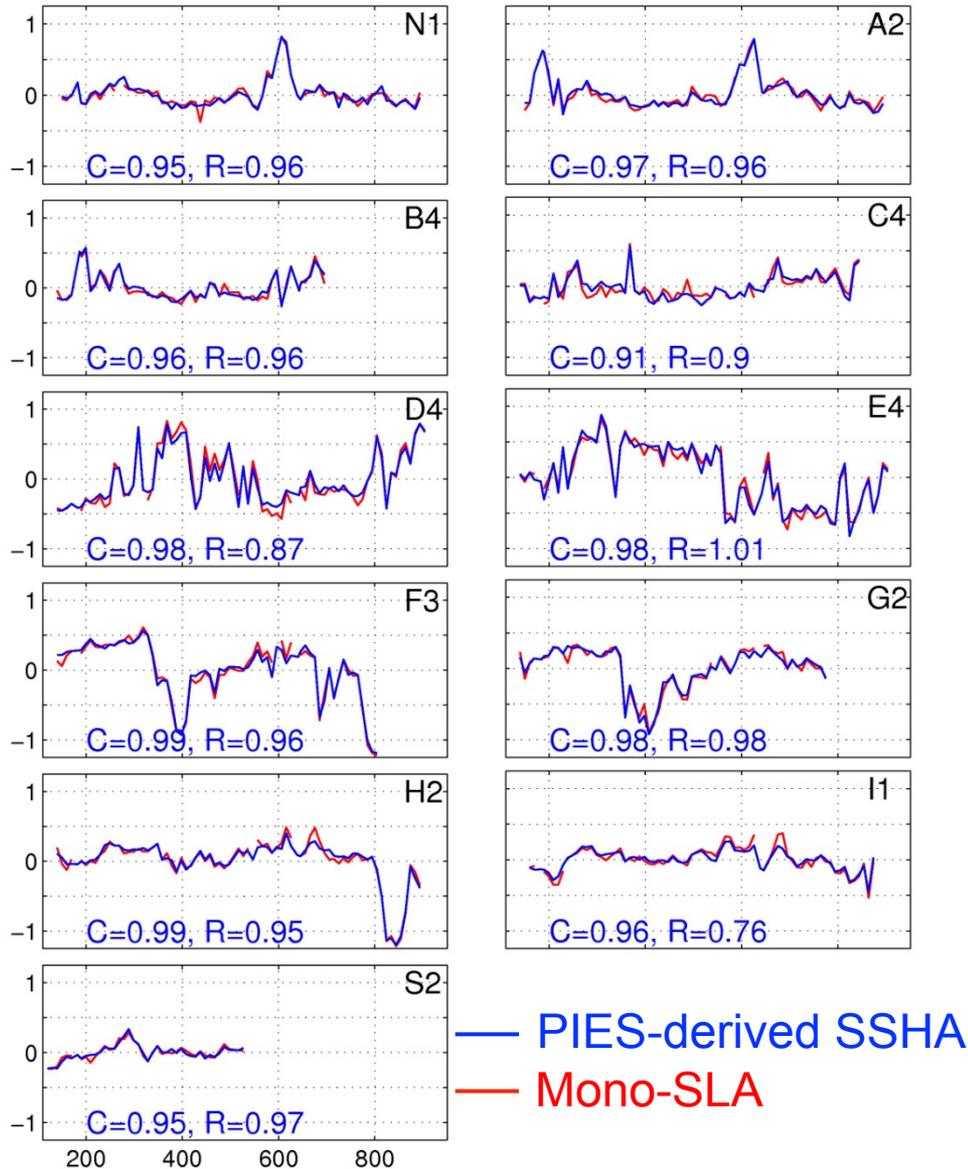
$\rho$  = Density

Typical sizes...

	$\Phi'/g$	$P'/\rho g$	$\eta'$	$\sigma_{\eta}$ (std dev)
	(m)	(m)	(m)	(m)
Kuroshio Jet	1.3	0.5	1.7	0.2
Quiet Intervals	0.3	0.2	0.4	0.1
Meanders, Rings	1.0	0.3	1.2	0.15

Steric and mass-load parts have similar amplitudes but different processes and different spectra and different lateral scales

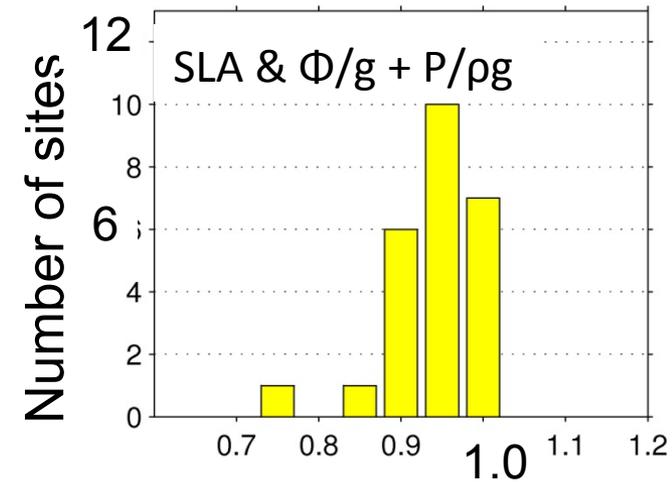
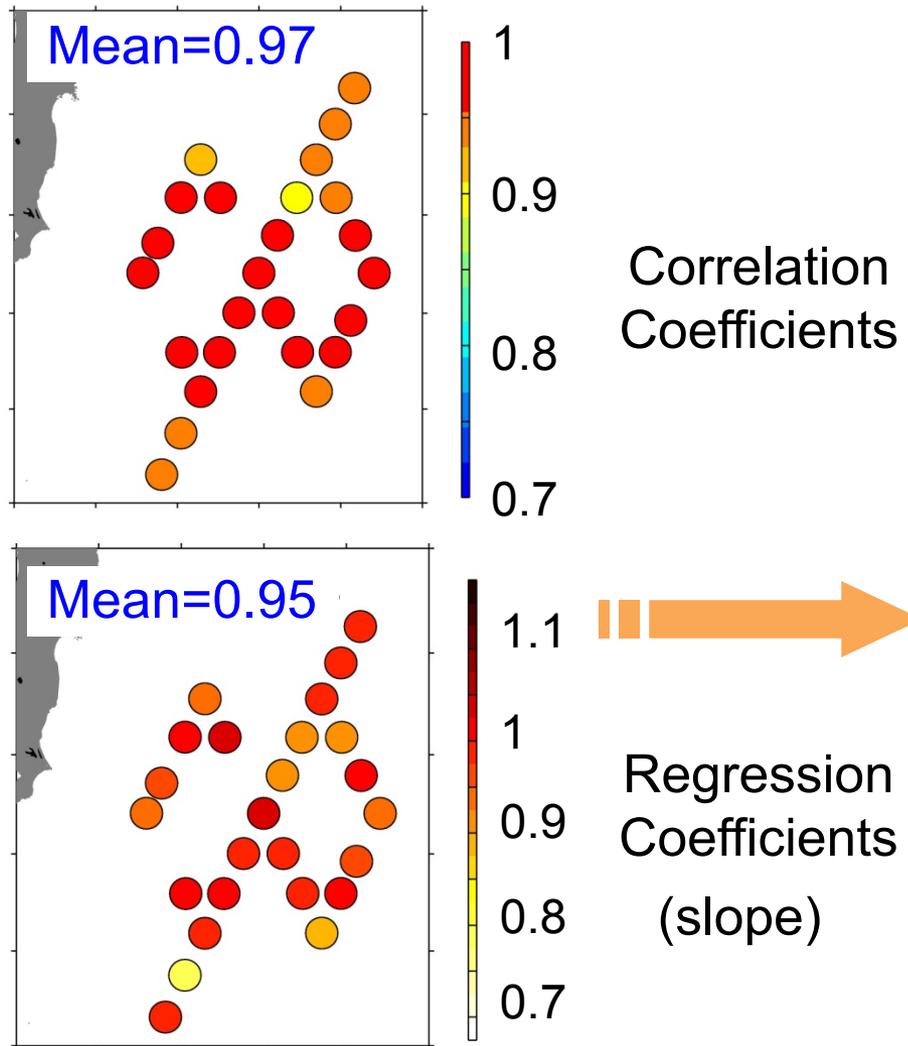
# Mono-SLA and PIES-derived SSHA



PIES data were interpolated to times when Jason-1 passed overhead.

Excellent agreement when both steric and mass-loading parts are included.

# Mono-SLA and PIES-derived SSHA



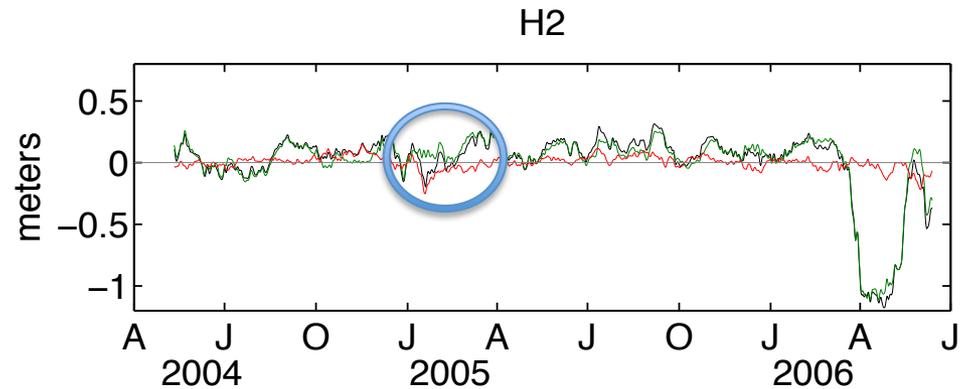
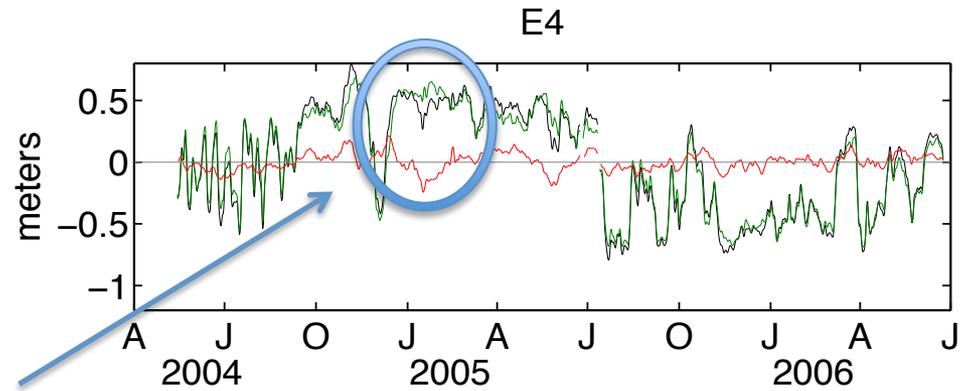
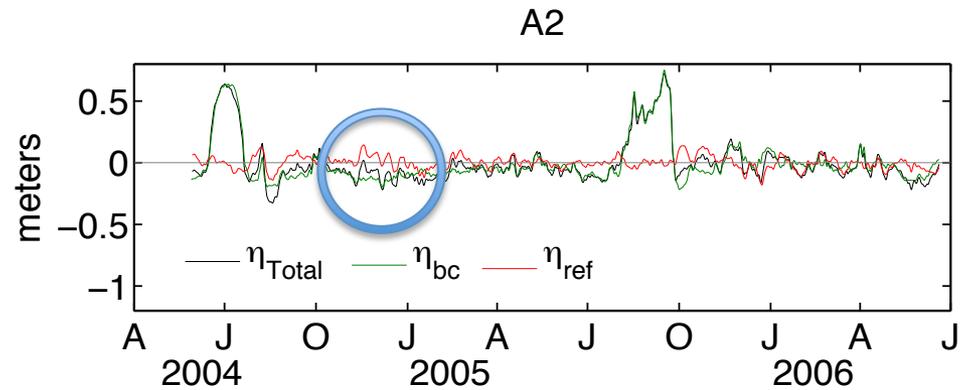
$$\eta_{\text{tot}} = \Phi'/g + P'/\rho g$$

$$= \eta_{bc} + \eta_{\text{ref}}$$

KESS northern, middle & southern sites

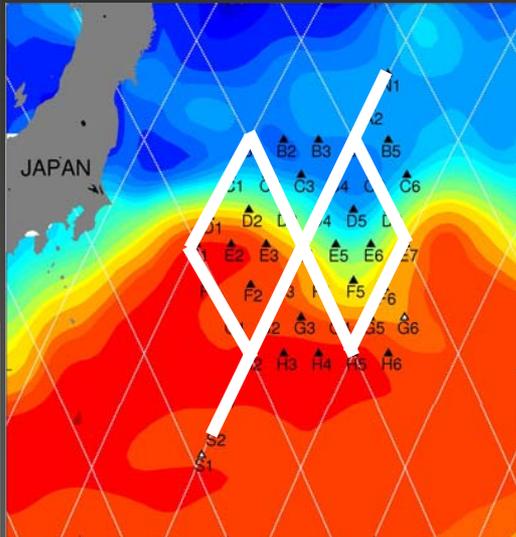
-steric part (bc) large

-mass-loading (ref) is important.



# Rms Errors (along-track 26 sites)

Based on Kuroshio and Agulhas & better for SWOT



❖ Observed rms differences  
Mono-SLA **6.8 cm**

❖ Predicted rms difference

$$\sqrt{(4.6 \text{ to } 5.1)^2 + (3.8 \text{ to } 8.7)^2 + 1.4^2} = \mathbf{6.1 \text{ to } 10.2 \text{ cm}}$$

when error caused by geographical position mismatch between Jason-1 and PIES  
= 1.4 cm (1 km offset)

❖ Error Budget for AVISO : **4.6-5.1 cm**

GDR Corrected SSH	3.3 cm
Post-processing - IB	2.5-3.0 cm
- Tides	2.0-2.5 cm

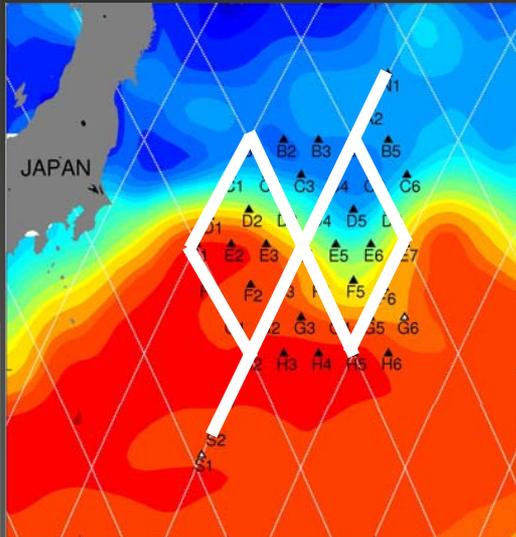
From Baker-Yeboah (2008)

❖ Error Budget for PIES : **3.8-8.7 cm**

Sea state scatter	0.2 cm
Sea state bias	0.1 cm
Tides	0.1 cm
Pressure drift	1.0 cm
Mooring motion	0.2 cm
Spline-curve Lookup	3.7-8.6 cm

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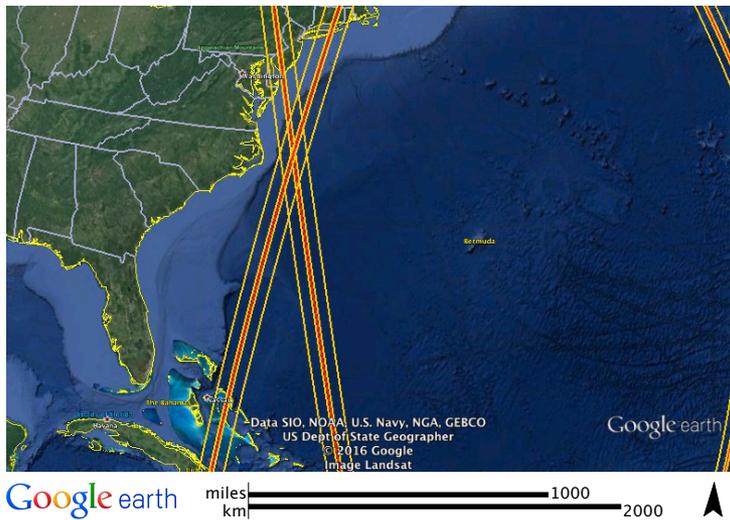
Sea state scatter	<del>0.2 cm</del>
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In order to improve the accuracy of independent SLA measurements, and to measure SLA on small 5-30 km lateral scales, where do we need more S,T,P measurements?

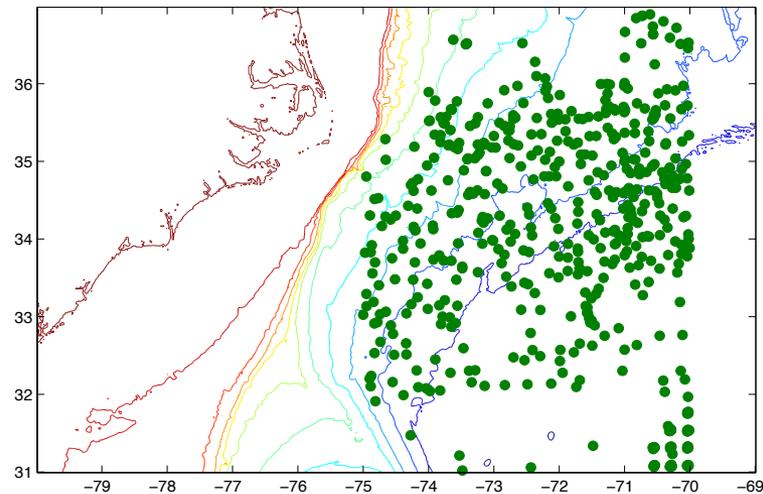
The main error contribution to  $SLA_{PIES}$  is from scatter in  $\Phi(\tau)$  lookup.

Examine next...

- Estimate  $\Phi_2'$  contributions from 2<sup>nd</sup> BC mode in full water column
- Bottom-intensified processes (TRW's, DWBC)
- Near-surface SQG processes and streamers (mixed layer to 150 m?)
- mid-column filaments (main pycnocline, 200 m – 700 m)



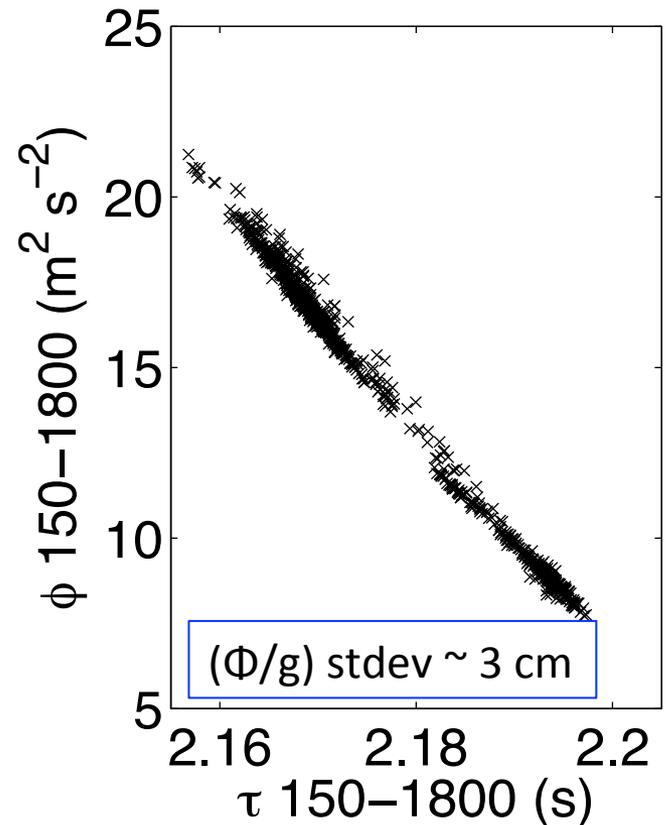
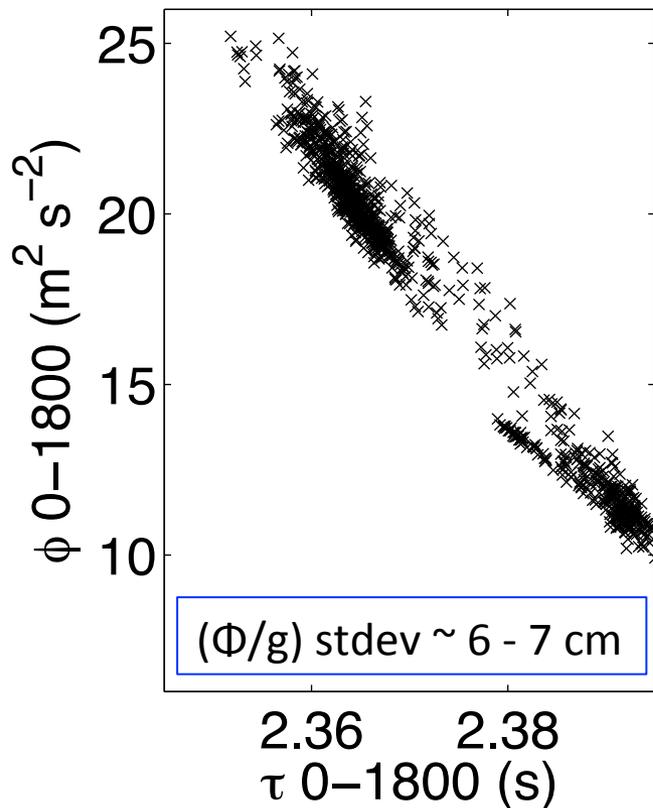
CTD profiles  
 (ship & Argo) →  
 for  
 $\Phi$  vs.  $\tau$  plots



Much of the  $\Phi(\tau)$  scatter comes from top 150m.

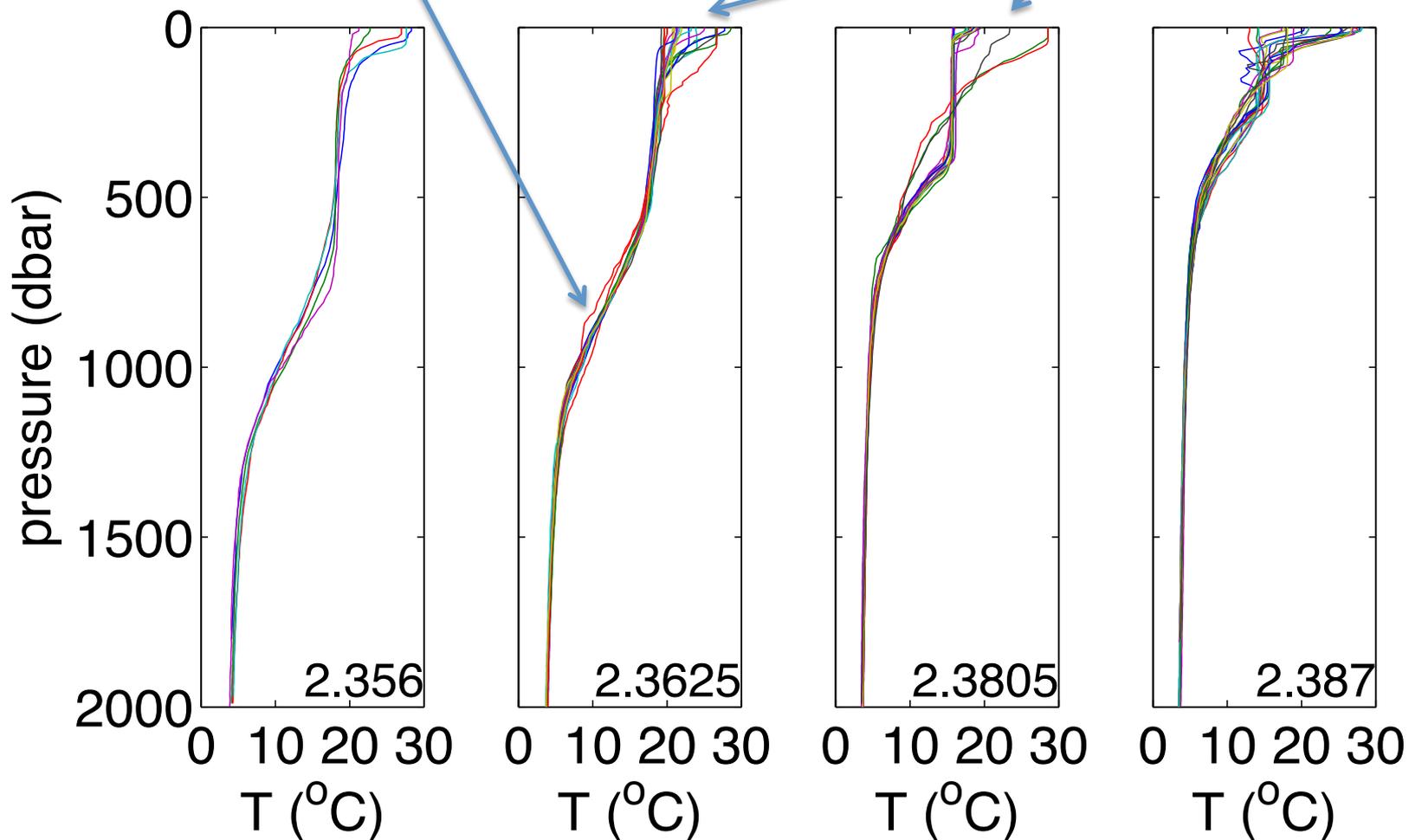
And much of that is due to seasonal warming, for which we can adjust.

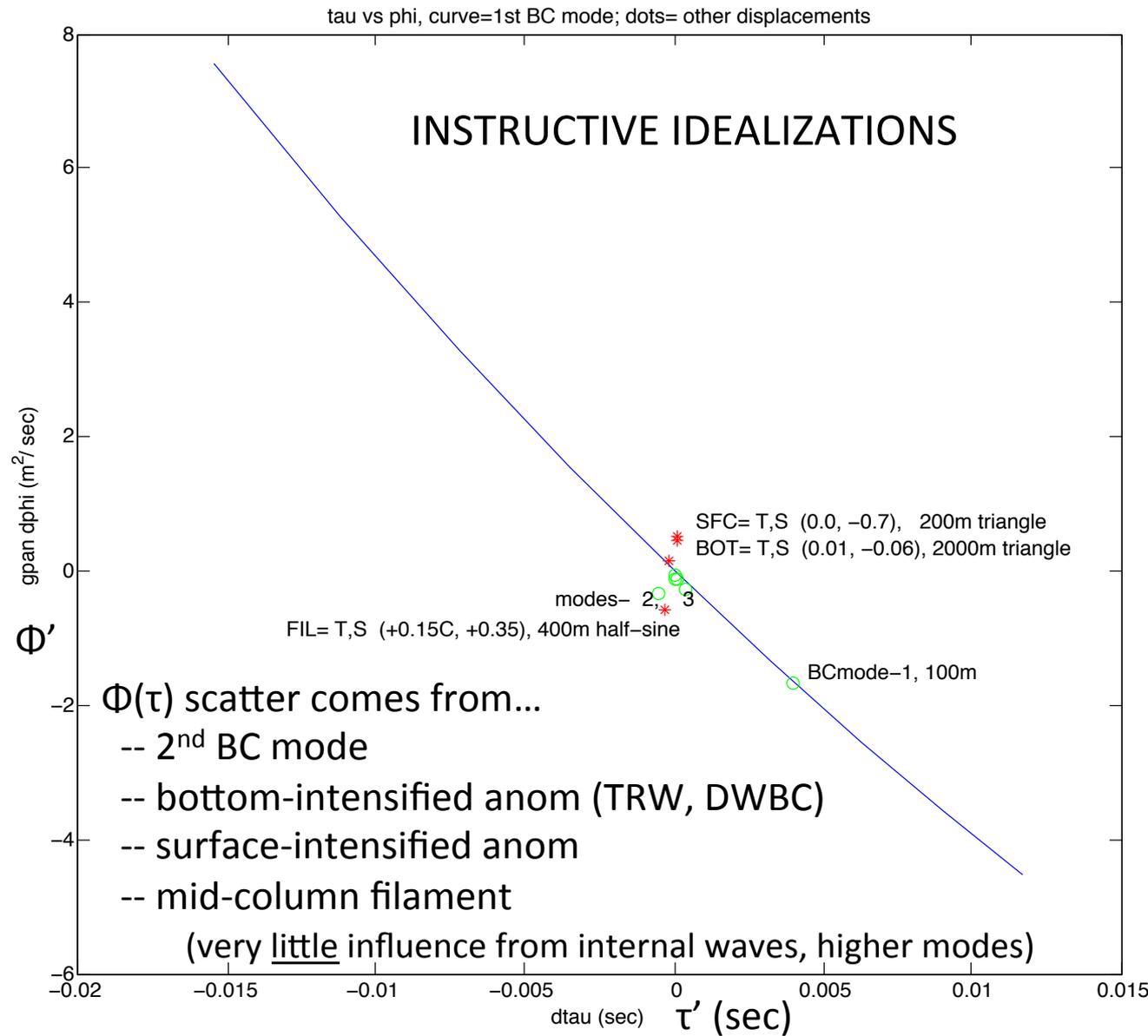
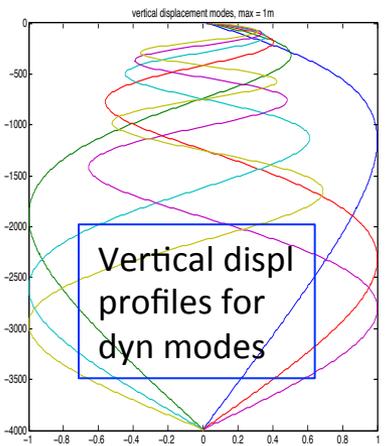
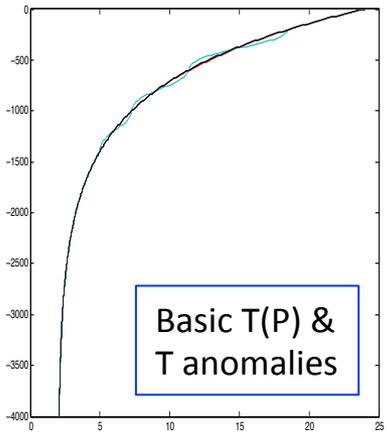
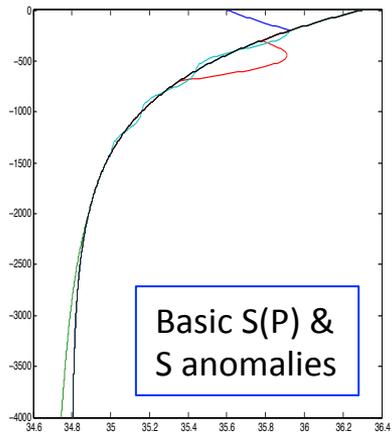
Less scatter than Agulhas & KESS.



For guidance about what causes scatter in  $\Phi(\tau)$  examine temperature profiles, grouped in narrow bins of  $\tau$  (1 ms).

These show subsurface intrusions (likely filaments) and surface warming





Combine CPIES with moorings, having many S, T, P recorders vertically, to determine  $S'$ ,  $T'$  anomalies relative to the GEM profile, will give the most accurate  $\Phi$  measurement.

Laterally space some CPIES closely ( $\sim 5\text{km}$ ).

Additional CPIES can be added to observe longer lateral scales.

Final points...

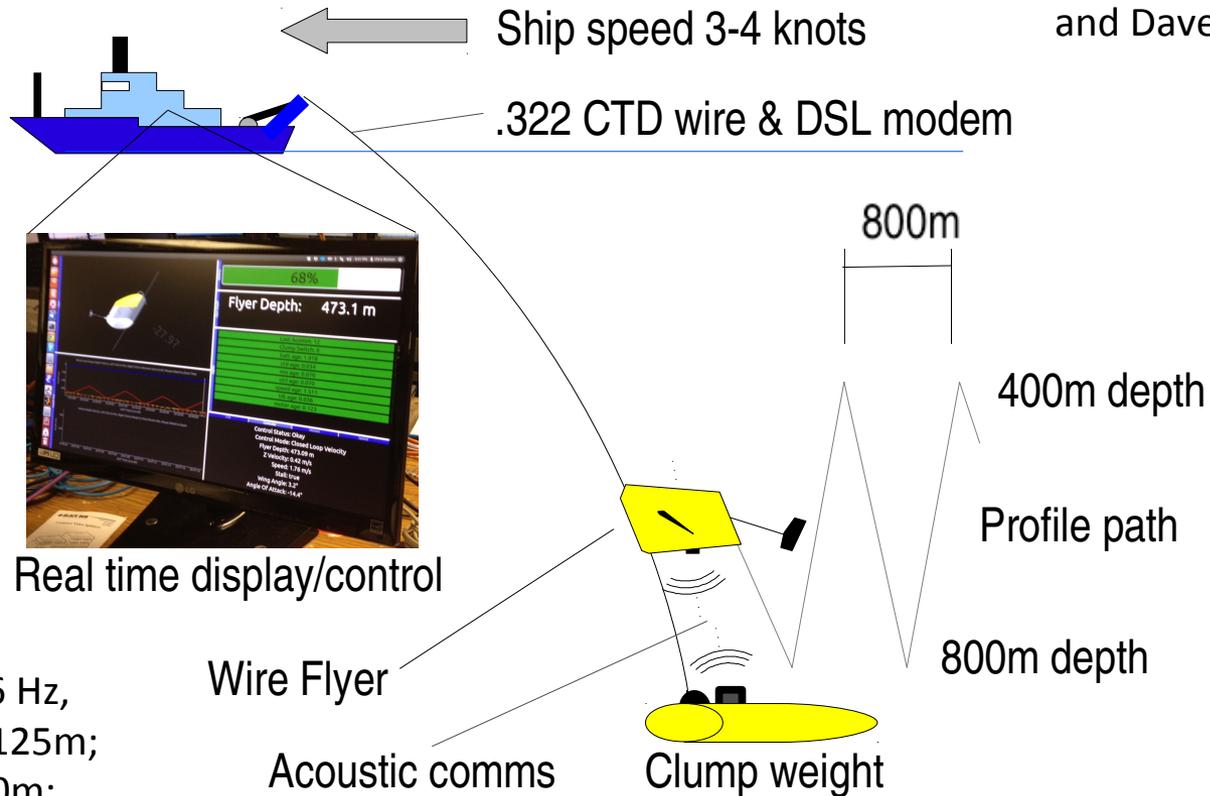
While the focus of this presentation is on understanding and reducing noise in full water column geopotential height  $\Phi'$  estimated from PIES  $\tau$ ,

1) remember also that bottom current meters are important for leveling and dedrifting the bottom  $P'$  measurements.

2) Wire-flyer = new fine lateral AND vertical scale capability combined

# WIRE FLYER

developed at URI by  
Chris Roman, Dave Hebert,  
and Dave Ullman



FastCat CTD 16 Hz,  
vertical res. 0.125m;  
horiz res. ~ 500m;  
transects ~ 178km

Shipboard surveys: *in situ* calibration CTD profiles for cPIES; with LADCP/ SADCP;  
estimate shear-to-strain ratios;  
unprecedented horizontal and vertical structure of streamers and filaments

END -- QUESTIONS?



# SWOT CalVal Planning Meeting

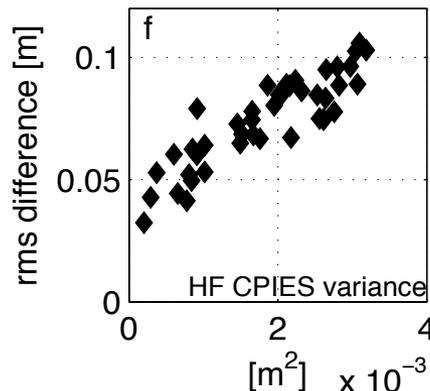
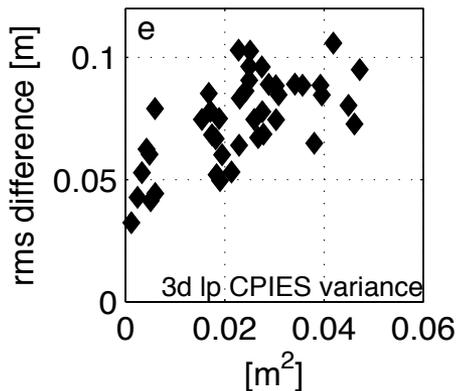
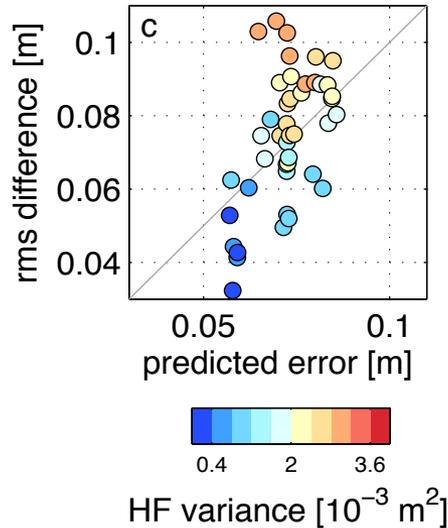
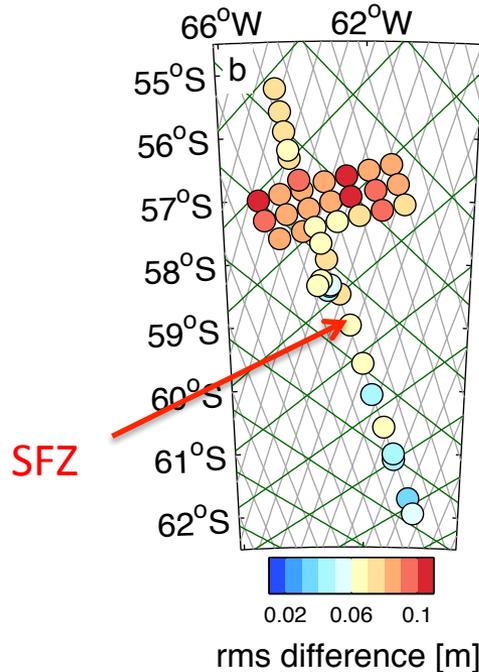
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- PIES intro;  $\tau$  and P,T
  - $(SSH' - IB') = \Phi'/g + P'/(g\rho)$
  - $\tau$  integral & geopotential  $\Phi$  integral; plot  $\Phi$  vs  $\tau$  with scatter; signal/noise ratio
  - Published examples of PIES measuring SSH, compared to altimeter SSH; uncertainties
- GOAL implied for SWOT calval:
  - Understand the measurement accuracy of SSH from SWOT, particularly on lateral scales from 5 km – 100 km
  - Compare against local independent trusted measurements of SSH
- Detail PIES\_SSH uncertainty; where do we need more STP measurements?
  - T(P) and S(P) profiles and wiggles affecting  $\tau'$  and  $\Phi'$  differently
  - Surface/ bottom / mid-column anomalies
  - Higher dynamical modes (2nd mode – implications near DWBC?)
  - Measurement error – P drift; CPIES geostrophic leveling (deploy 6+ mos in advance?)
- Implied/ deduced lateral scales of features

# Drake Passage (cDrake experiment)

AVISO MSLA predicted mapping and CPIES errors agree well with observed rms differences.



Much of the  $|\eta_{\text{alt}} - \eta_{\text{PIES}}|$  difference in excess of 0.05 m arises from high frequency <20 d variance that was unresolved by the Jason 10 d sampling.

# PIES

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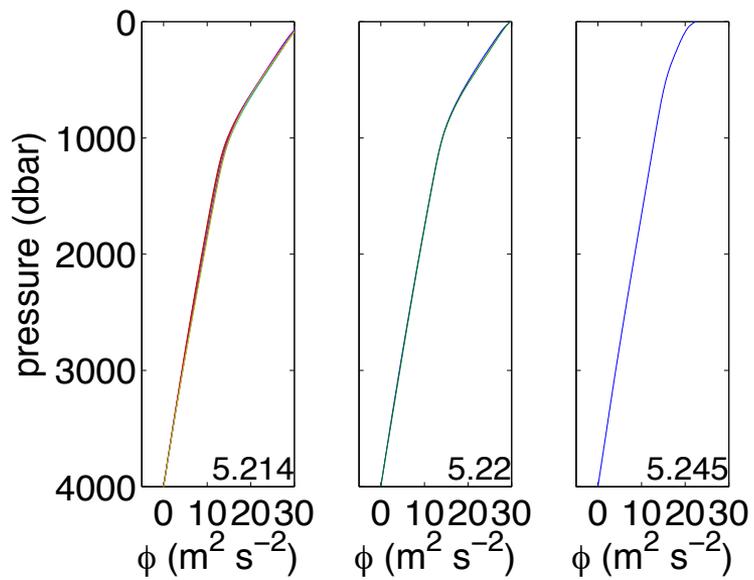
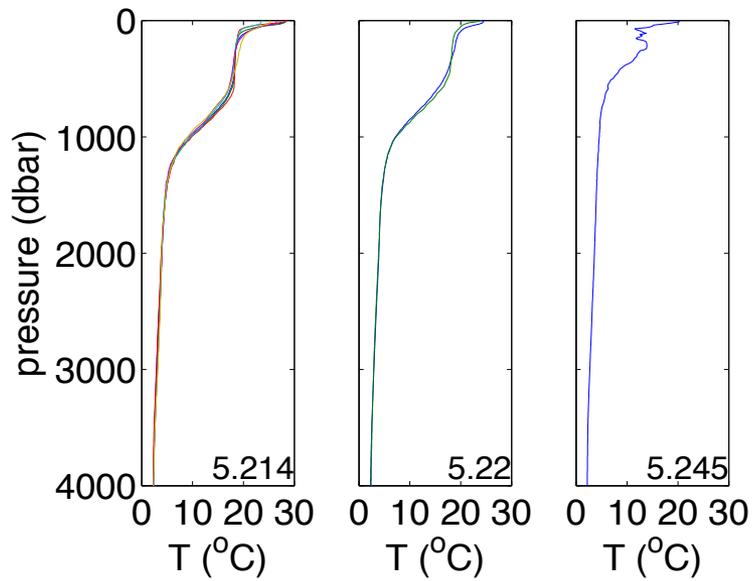


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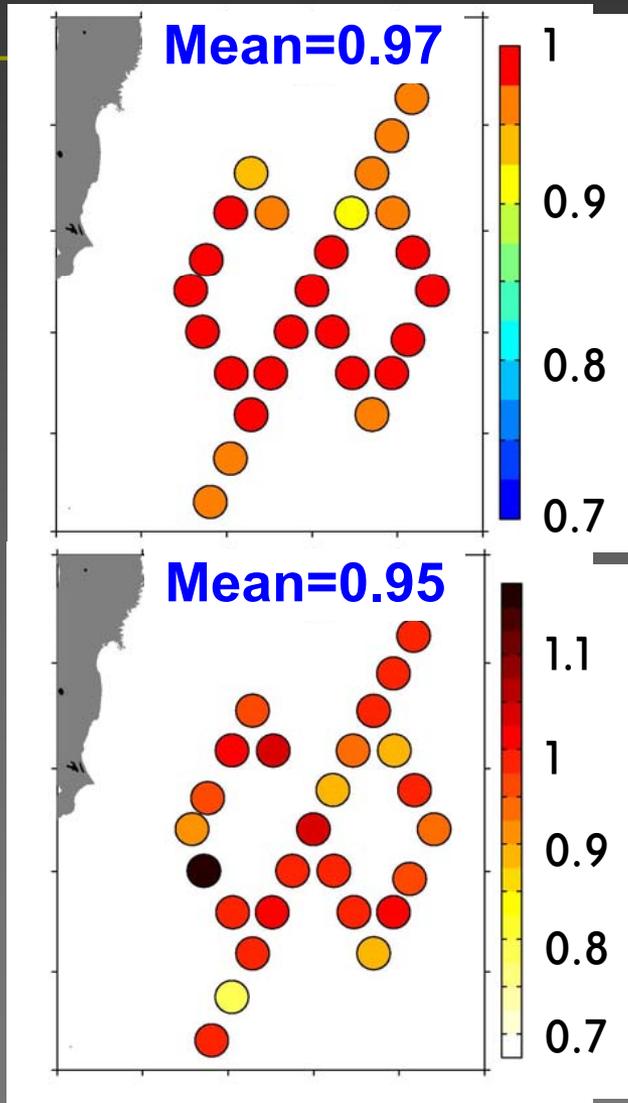
$$\tau = \int (1/c) dz$$

$$\Phi = \int (1/\rho) dp$$

→ geopotential  $\Phi$



# Mono-SLA and PIES-derived SSHA



**Correlation  
Coefficients**

**Regression  
Coefficients  
(slope)**

